

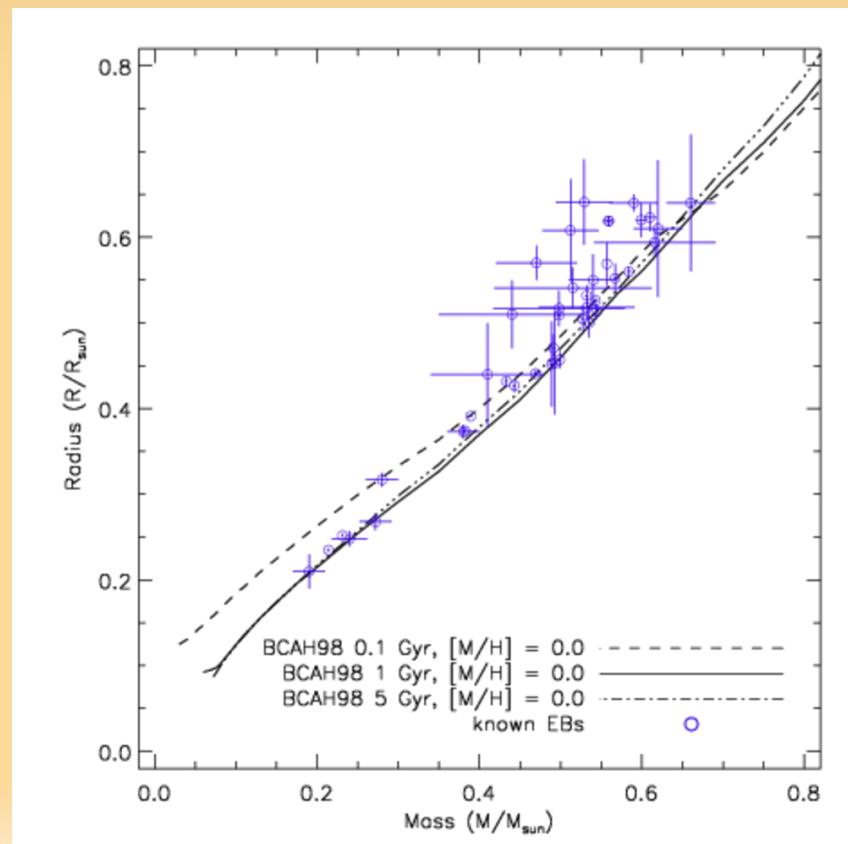
Testing Low Mass Stellar Models with M-dwarf Eclipsing Binaries from SDSS Stripe 82

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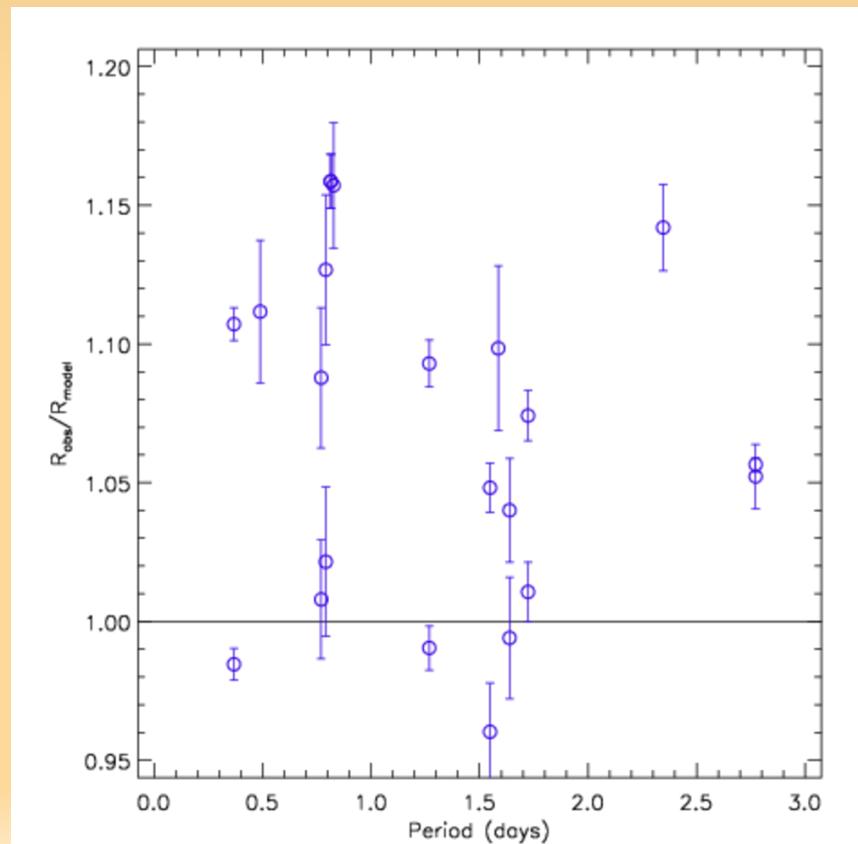
Low Mass Stars and Stellar Models

- Low mass stars in eclipsing binaries allow direct measurement of stellar properties
 - Mass, radius, effective temperatures, luminosities, chemical composition
- Observed stellar radii are $\sim 10 - 15\%$ larger than modeled radii
- Are tidally induced magnetic fields in close binaries the cause of this discrepancy?

Lower MS Mass – Radius Relation



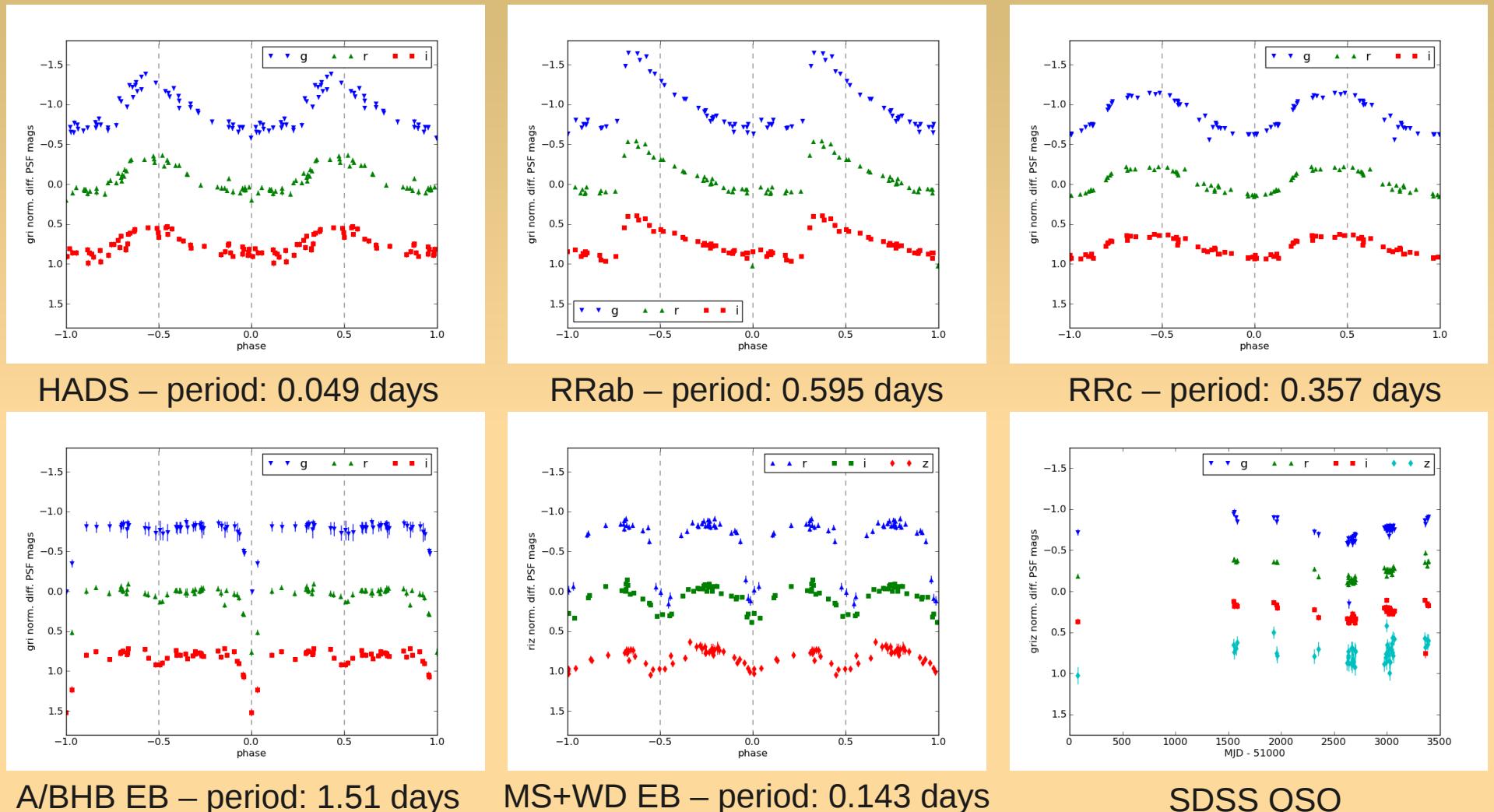
dM EB Orbital Period – Radius Discrepancy



Outline

- Variability in SDSS Stripe 82
- Eclipsing binary systems from this dataset
 - Modeling
 - Parameter distributions
 - Follow-up candidate selection
- Follow-up observations
 - SDSS J2122 – dM0 + dM1 EB
 - SDSS J0211 – dM1 + dM2 EB
- Future work

Variables from SDSS Stripe 82



- ~ **2 million** total point sources identified
- ~ **1.3 million** point sources with:
 - ≥ 10 detections, SDSS $r < 22.0$
- **16796** total variable point sources
- **781** periodic variables
 - **382** EBs
 - **288** RR Lyrae, **49** HADS
 - ~ **60** other (rotation, CV, ELL, etc.)



Reset form

Search catalog

Output type

 HTML table CSV text

Sort output by

- PSF r mag (ascending) RA (ascending) Stetson variability index (ascending) Object tags (ascending) Periodic variable type (ascending)
 PSF r mag (descending) RA (descending) Stetson variability index (descending) Object tags (descending) Periodic variable type (descending)

Object name constraints

 no object name constraints

separate names by commas
use the form JXXXXXX.XX+/-XXXXXX.X

 search by SDSS J name(s)

separate names by commas
22982, 140427, 36, 30313, 25412

 search by object MB name(s)

Position constraints

 no position constraints

use decimal degrees (J2000 coords)

min α max α min δ max δ

use decimal degrees (J2000 coords)

α and δ using a radius of arcmin

Magnitude constraints

 use PSF u mag limitsmin to max mag use PSF g mag limitsmin to max mag use PSF r mag limitsmin to max mag use PSF i mag limitsmin to max mag use PSF z mag limitsmin to max mag

Color constraints

- use PSF u-g color limits min to max mag
 use PSF g-r color limits min to max mag
 use PSF r-i color limits min to max mag
 use PSF i-z color limits min to max mag

Object tag constraints

- objects are tagged as **ANY** of the following...
- | | | |
|---|--|---|
| <input type="checkbox"/> SDSS QSO | <input type="checkbox"/> QSO color candidate | <input type="checkbox"/> SDSS white dwarf |
| <input type="checkbox"/> SDSS hot subdwarf | <input type="checkbox"/> SDSS-I variable | <input type="checkbox"/> SDSS standard |
| <input type="checkbox"/> RR Lyrae candidate | <input type="checkbox"/> white dwarf/sdO/sdB | <input type="checkbox"/> A/BHB |
| <input type="checkbox"/> F-turnoff/sdF | <input type="checkbox"/> low-metallicity | <input type="checkbox"/> F/G |
| <input type="checkbox"/> G dwarf | <input type="checkbox"/> K giant | <input type="checkbox"/> AGB |
| <input type="checkbox"/> K dwarf | <input type="checkbox"/> sdM | <input type="checkbox"/> M dwarf |
| <input type="checkbox"/> MS+WD pair | <input type="checkbox"/> brown dwarf | <input type="checkbox"/> unknown |
- objects are tagged as **ALL** of the following...
- objects are tagged as **NONE** of the following...

Periodic variability constraints

- no periodic variability constraints
 object(s) are not tagged as periodic variables
 object(s) are tagged as periodic variables of any type
 object(s) are tagged as periodic variables of specific type(s):

- | | |
|---|---|
| <input type="checkbox"/> eclipsing binary candidate | <input type="checkbox"/> high amplitude Delta Scuti |
| <input type="checkbox"/> RRab RR Lyrae | <input type="checkbox"/> RRc RR Lyrae |
| <input type="checkbox"/> ellipsoidal binary | <input type="checkbox"/> spot rotation |
| <input type="checkbox"/> cataclysmic variable | <input type="checkbox"/> unknown |
- look for periodic variables with secure periods only
 restrict periods to values min to max days

Firefox ▾ Search results

shrike.pha.jhu.edu/cat-search/do-var-search

Google



Search the light-curve catalogs > Search the variable point source catalog > Search results

Your query was translated to the following SQL statement:

```
SELECT a.mb as mb, a.name as name, a.nobs as nobs, a.ra as ra, a.dec as dec, a.star_tags as star_tags, a.sdss_psfs as sdss_psfs, a.sdss_psfg as
sdss_psfg, a.sdss_psfr as sdss_psfr, a.sdss_psfi as sdss_psfi, a.sdss_ps fz as sdss_ps fz, a.lc_csv_fname as lc_csv_fname, a.lc_pdf_fname as
lc_pdf_fname, b.vartype as vartype, b.html_desc_fname as html_desc_fname, a.stet_ri as stet_ri, a.uid, b.period FROM probable_variables a LEFT OUTER
JOIN periodic_variables b ON (a.uid = b.uid) WHERE (((a.star_tags like %K dwarf%) or (a.star_tags like %M0%) or (a.star_tags like %M1%) or
(a.star_tags like %M2%) or (a.star_tags like %M3%) or (a.star_tags like %M4%) or (a.star_tags like %M5%) or (a.star_tags like %M6%) or (a.star_tags
like %M7%) or (a.star_tags like %M8%) or (a.star_tags like %M9%)) and ((a.star_tags not like %low-metallicity%) and (a.star_tags not like %K
giant%))) and ((b.vartype like %EB%) AND (b.vartype not like %uncertain%)) ORDER by a.stet_ri desc;
```

Notes:

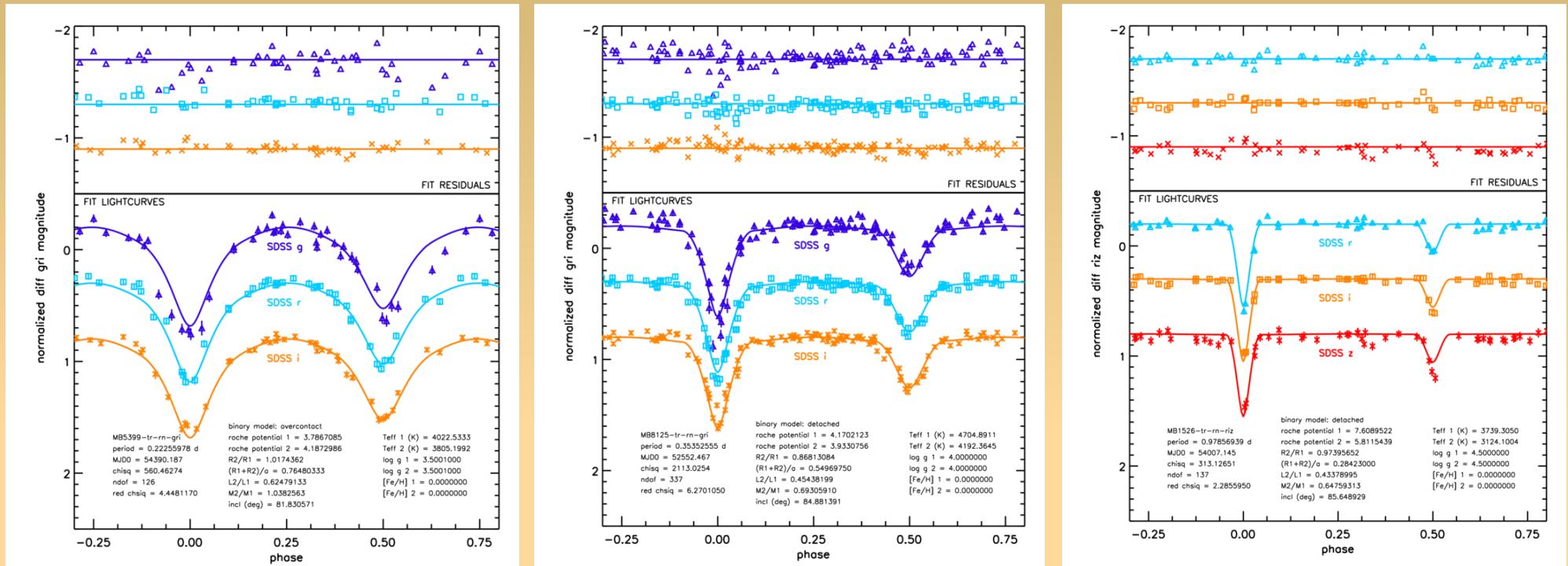
- Objects in the table below are sorted by decreasing Stetson variability index.
- Clicking on the cutout image will take you to the SDSS DR7 Object Explorer page for this object. If it is not present in DR7, then the object returned might not be the one you were looking for. This is because the Object Explorer searches a rather large 5 arcmin radius around the coordinates given.
- Detected periodic variable objects are noted by type below and links to their description pages are provided.
- The **LC CSV** column links to a tarball containing CSVs of the light-curves of the object. The **LC info** column links to a page with light-curves of the object.

Run another search query.

The following 59 object(s) match your search criteria:

object	RA [deg]	DEC [deg]	nobs	object tags	u	g	r	i	z	LC CSV	LC info	SDSS cutout
MB58833 SDSS J205650.83-005633.5 periodic var: dK EB period: 0.274258 days	314.212	-0.943	45	K dwarf / SDSS-I variable	21.15	19.40	18.69	18.50	18.31	get	view	
MB178470 SDSS J205122.50-010555.0 periodic var: dK EB period: 0.263172 days	312.844	-1.099	36	K dwarf / SDSS-I variable	19.87	18.30	17.72	17.53	17.42	get	view	
MB2162												

Stripe 82 Eclipsing Binaries



gK EB – period: 0.223 days

dK EB – period: 0.353 days

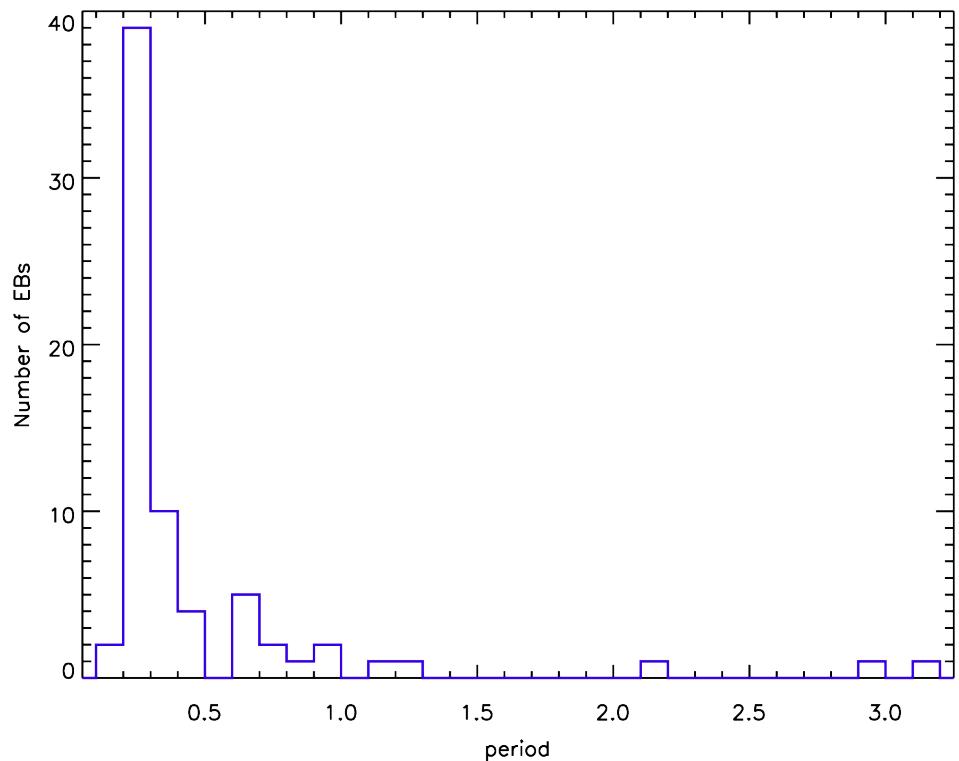
dM EB – period: 0.979 days

- **382 total EBs**
 - **162** EBs with secure periods
 - **104** EBs suitable for LC modeling
- Use Wilson-Devinney model code
 - Fit SDSS *gri* LCs for $r-i < 0.7$
 - Fit SDSS *riz* LCs for $r-i > 0.7$

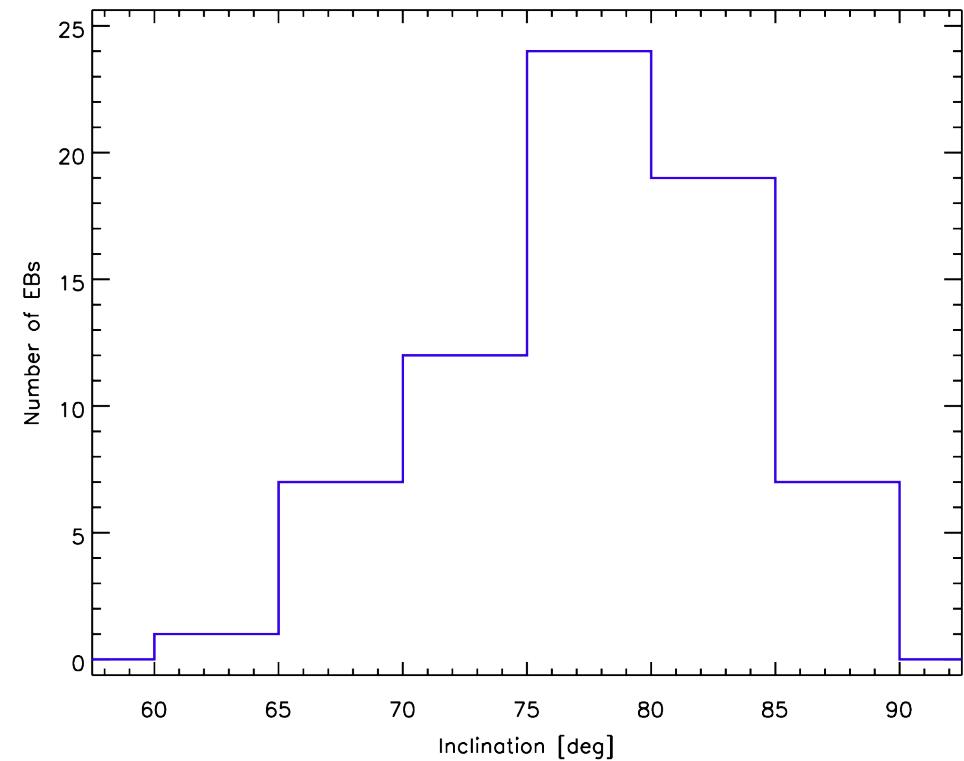
- Fix:
 - $e = 0$, no third light, no spots
 - $\log g$, $[Fe/H]$
- Fit:
 - R_2/R_1 , $(R_1+R_2)/a$
 - L_2/L_1 , M_2/M_1
 - inclination
 - MJD0
 - effective temperatures

EB Orbital Parameter Distributions

Orbital Period

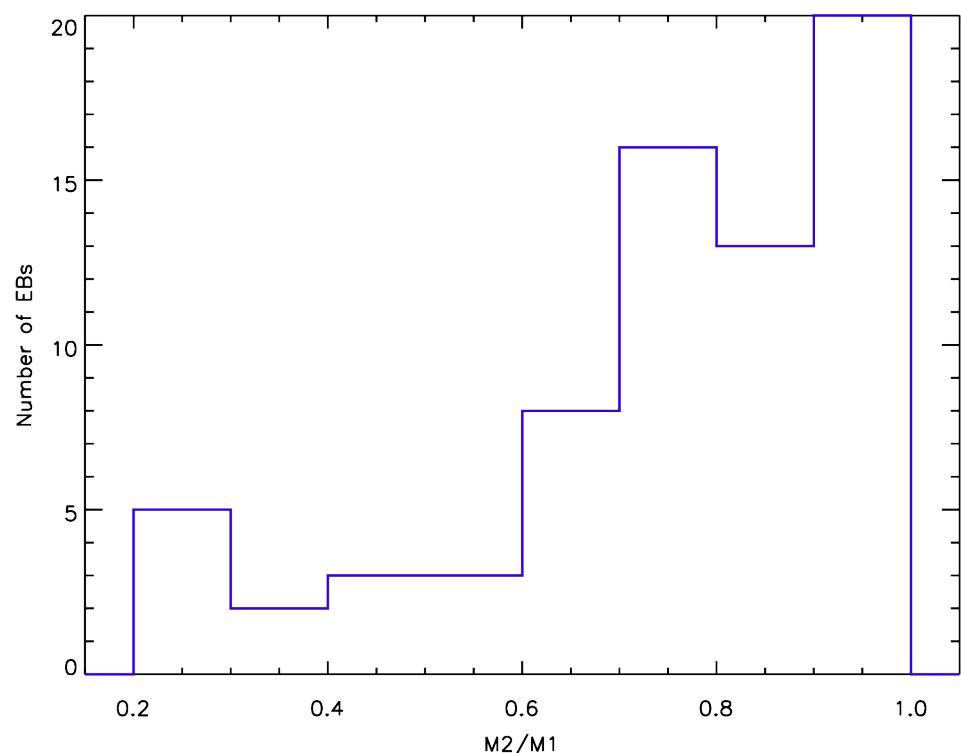


Orbital Inclination

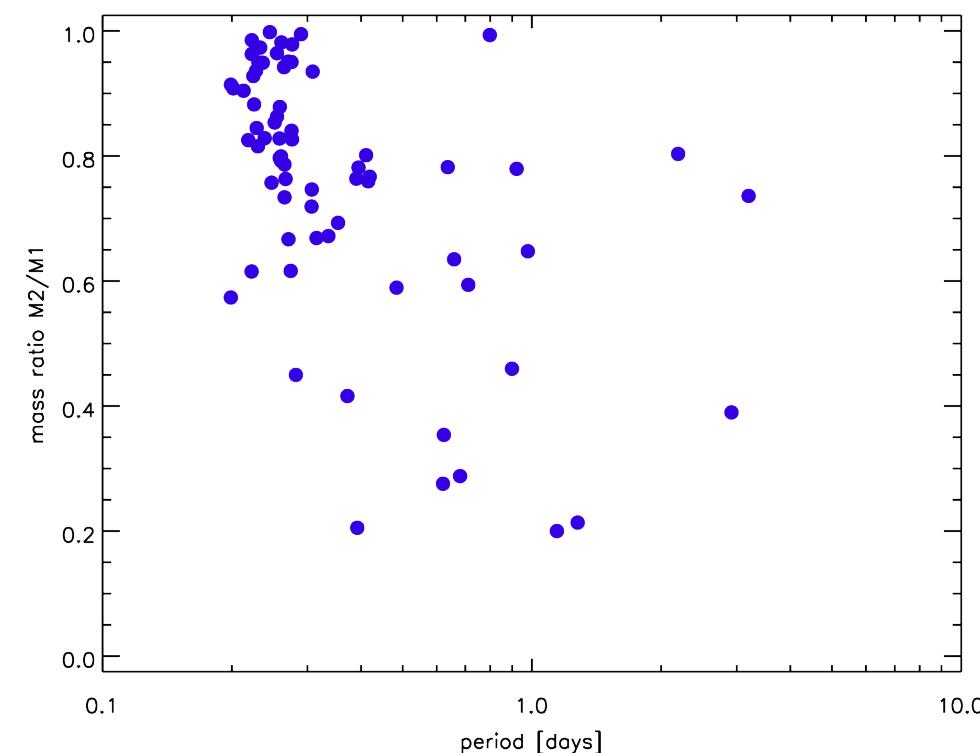


EB Stellar Parameter Distributions

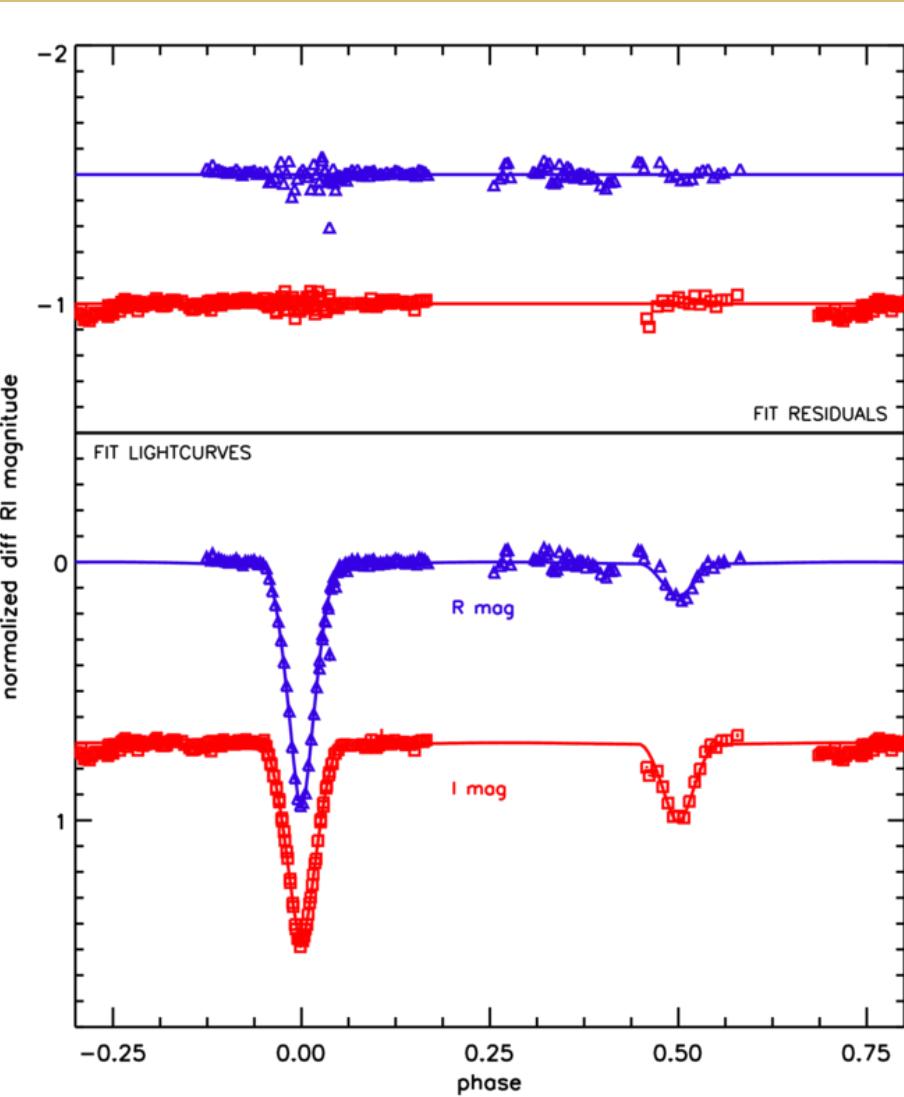
Photometric mass ratio M2/M1



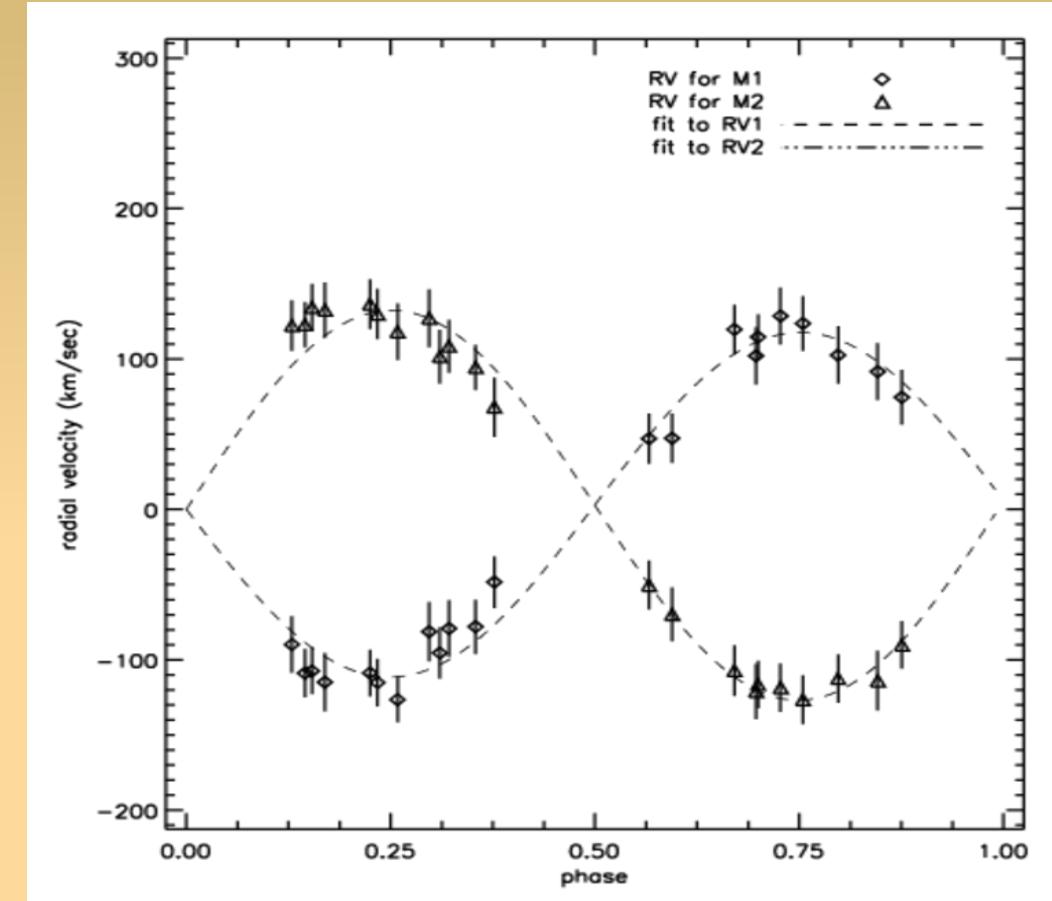
Photometric mass ratio vs. period



Two M-dwarf EBs – SDSS J2122



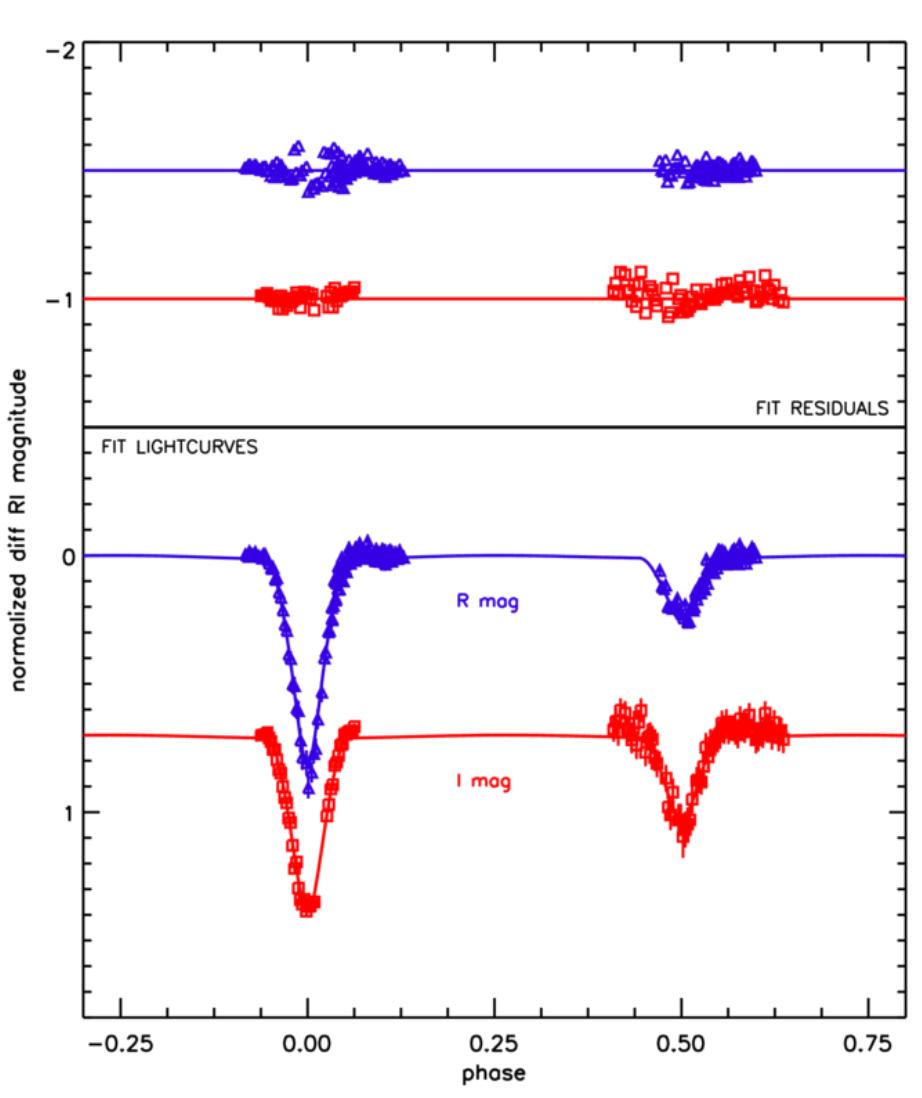
SDSS J212203.12-010053.3
 Period: 0.791032 ± 0.000147 days
 HJD0: 2455775.9743



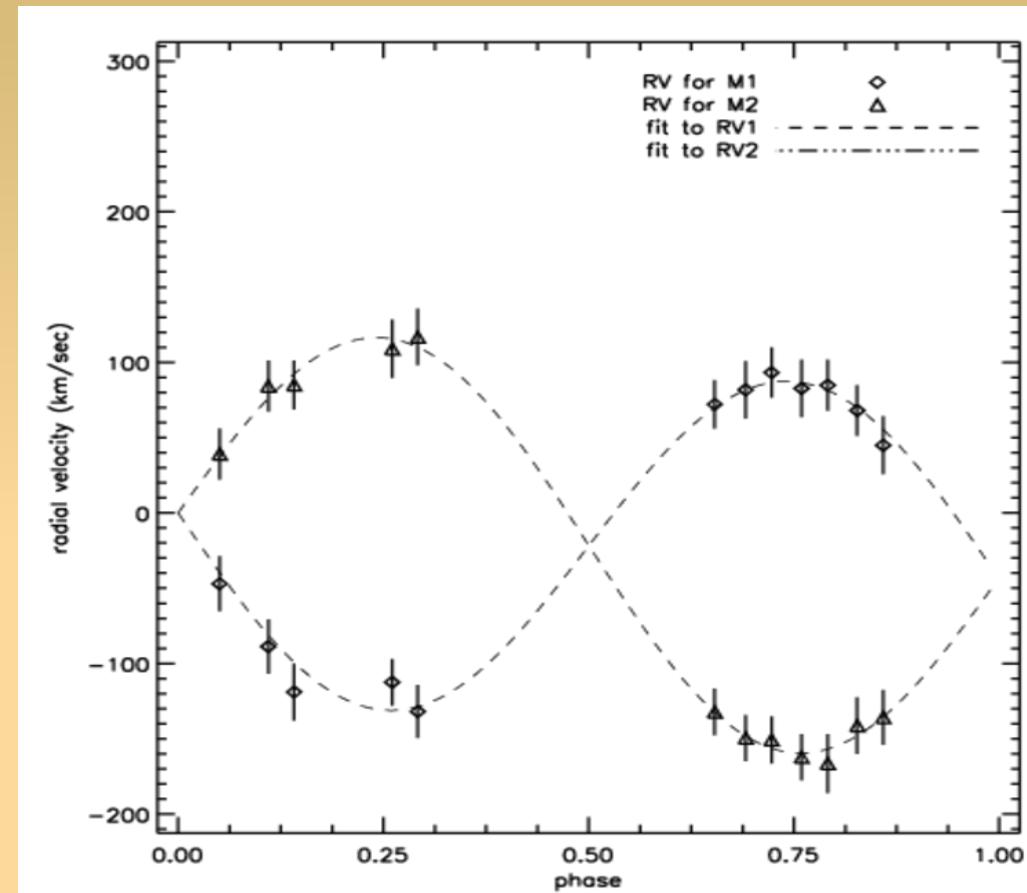
$$\begin{aligned} M_1 &: 0.616 \pm 0.075 M_{\odot} & R_1 &: 0.594 \pm 0.016 R_{\odot} \\ M_2 &: 0.532 \pm 0.059 M_{\odot} & R_2 &: 0.518 \pm 0.014 R_{\odot} \end{aligned}$$

$$\begin{aligned} T_{\text{eff},1} &: 3330 \pm 338 \text{ K, SpT: dM0} \\ T_{\text{eff},2} &: 2934 \pm 510 \text{ K, SpT: dM1} \end{aligned}$$

Two M-dwarf EBs – SDSS J0211



SDSS J021121.55-003808.3
Period: 0.623655 ± 0.000406 days
HJD0: 2455879.7346

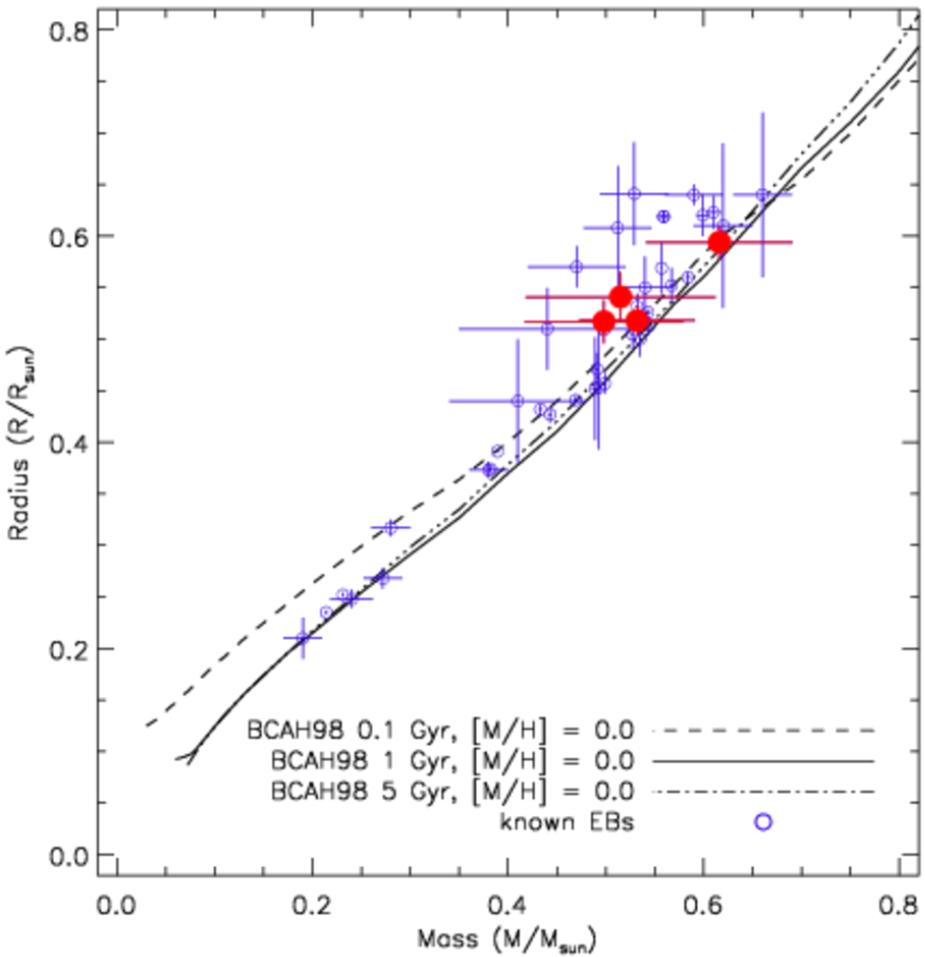


$$\begin{aligned} M_1 &: 0.515 \pm 0.097 M_{\odot} & R_1 &: 0.541 \pm 0.024 R_{\odot} \\ M_2 &: 0.498 \pm 0.082 M_{\odot} & R_2 &: 0.517 \pm 0.021 R_{\odot} \end{aligned}$$

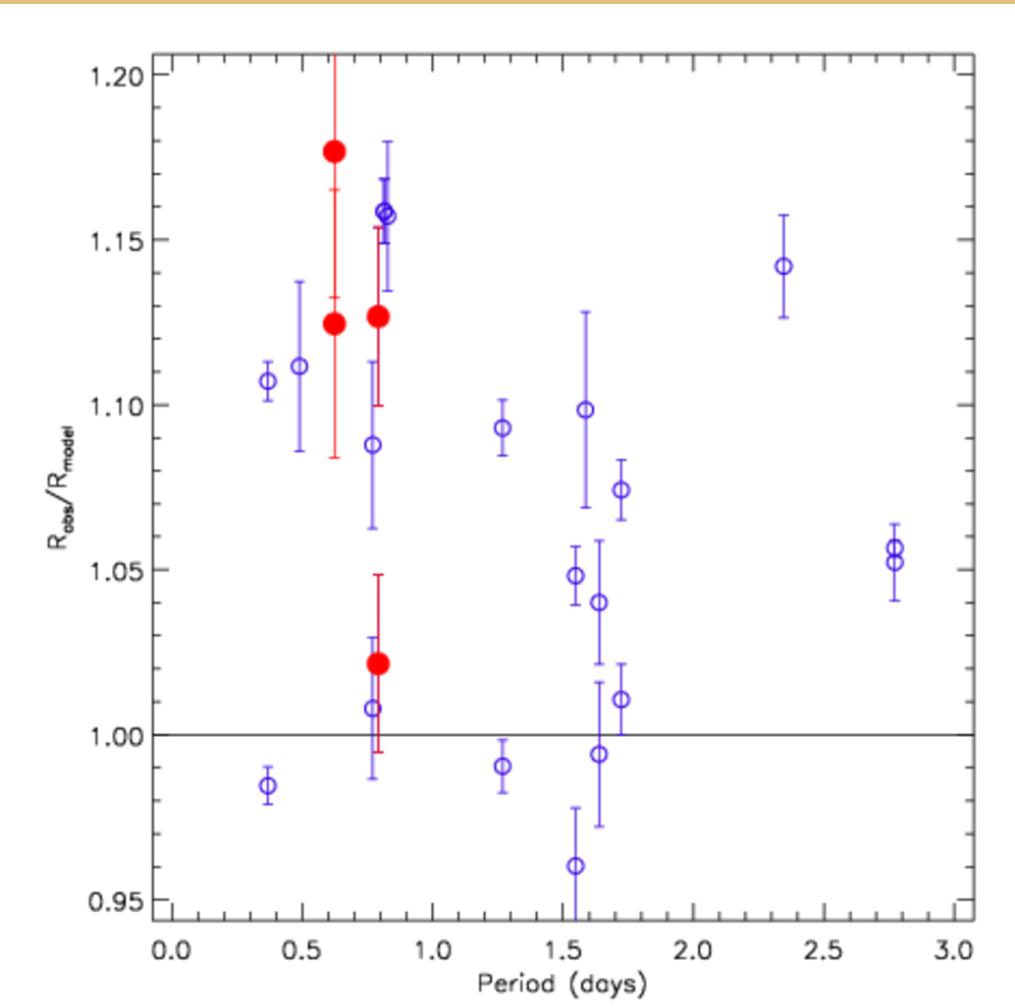
$$\begin{aligned} T_{\text{eff},1} &: 3192 \pm 473 \text{ K, SpT: dM1} \\ T_{\text{eff},2} &: 2836 \pm 619 \text{ K, SpT: dM2} \end{aligned}$$

Low Mass Stellar Models Revisited

Lower MS Mass – Radius Relation



dM EB Orbital Period – Radius Discrepancy



Future Work

- More EB catalog science
- Estimate absolute masses and radii of Stripe 82 dM EB sample directly
 - cf. Coughlin et al. 2011 (Kepler sample of EBs)
- More follow-up light-curves and RVs for final target EBs
 - Observations this week
- Improve modeling of final two EB targets
 - Better mass estimates needed

Summary

- ~1.3 million point sources extracted from SDSS Stripe 82
- 16796 variable point sources
 - 781 periodic variable stars
 - 382 EBs, 288 RR Lyrae, 49 HADS, ~60 other
 - Search: <http://shrike.pha.jhu.edu/cat-search/var-search>
 - Browse: <http://shrike.pha.jhu.edu/stripe82-variables>
- 2 M-dwarf EBs with follow-up observations to get M + R
 - SDSS J2202 (dM0 + dM1):
 - $M_1: 0.616 \pm 0.075 M_{\odot}$, $R_1: 0.594 \pm 0.016 R_{\odot}$
 - $M_2: 0.532 \pm 0.059 M_{\odot}$, $R_2: 0.518 \pm 0.014 R_{\odot}$
 - SDSS J0211 (dM1 + dM2):
 - $M_1: 0.515 \pm 0.097 M_{\odot}$, $R_1: 0.541 \pm 0.024 R_{\odot}$
 - $M_2: 0.498 \pm 0.082 M_{\odot}$, $R_2: 0.517 \pm 0.021 R_{\odot}$
- Observed radii are larger than predicted radii from models